

Calculating the mass of the Sun

Key Stage 4

Topics covered: Newton's law of gravitation, Kepler's third law

Watch the video "How do we know how old the Sun is?" <u>https://vimeo.com/88978362</u>



Kepler's third law states that the orbital period of a planet, T, squared is proportional to its distance from the Sun, R, cubed:

$$T^2 \propto R^3$$
 (1)

Newton's law of gravity states that the mutual gravitational force that the Sun and an orbiting planet feel is proportional to the mass of the Sun, M; the mass of the planet, m and is inversely proportional to the distance between the two squared, R^2 :

$$F = \frac{GMm}{R^2} \tag{2}$$

Kepler's third law can be derived from Newton's law of gravity, the result is this:

$$\frac{T^2}{R^3}M = \frac{4\pi^2}{G}$$
 (3)

T is in seconds, R is in kilometres, M is in kilograms and G is the gravitational constant = $6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-1}$.

If T is in years and R is in astronomical units (AU) then $T^2 = R^3$.



- 1. The orbital period of Saturn is 29.5 years. Calculate the orbital distance of Saturn in AU. Convert this into kilometres.
- 2. Convert the orbital period of Saturn into seconds.
- 3. Rearrange equation 3 and calculate the mass of the Sun.



Calculating the mass of the Sun: ANSWERS

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- 1. 9.54 AU; 1.43 x 10⁹ km
- 2. 930949200 seconds
- 3. The mass of the Sun is 2×10^{30} kg